## Hyper-Kamiokande

Federico Nova

**CPNR 2025** 

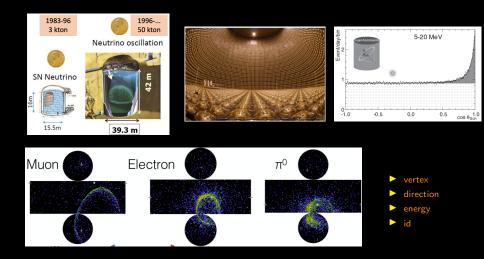


#### Outline:

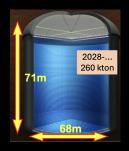
- 1. detector
- 2. physics program



## Kamiokande family

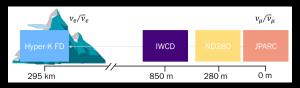


## Hyper-Kamiokande detector

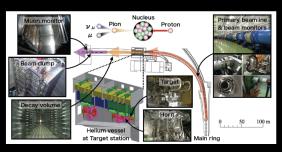


	SuperKamiokande	Hyper-Kamiokande	
site	Mozumi	Tochibora	
(rock overburden) ID PMTs	$^{(1\ km)}$ $\sim$ $11 k$	(600 m) ∼20k	
photo-coverage	(50 cm) 40%	(50 cm) 20%	
OD PMTs	2k	$ imes$ 2 single p.e. efficiency $3.6 \frac{k}{}$	
mass	(20 cm) <b>50 <mark>kton</mark></b>	(8 cm) 260 <mark>kton</mark>	
(fiducial mass)	(22.5 kton)	(188 kton)	

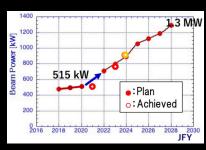
- ▶ 8× SK fiducial volume
- improved photo-sensors
- begin data taking 2028



#### J-PARC neutrino beamline

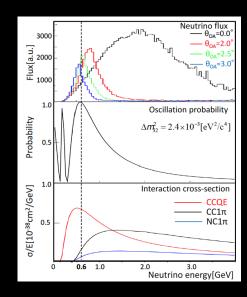


- ightharpoonup narrow-band neutrino beam ( $E \approx 600 \ \mathrm{MeV}$ )
- ▶ beam power to reach 1.3 MW (T2K ran stably at 800 kV in 2024)
  - upgrade main ring RF system:  $2.6 
    ightarrow 3.3 \cdot 10^{14} \; {
    m protons/pulse}$
  - improved horns power:  $250 \rightarrow 320 \text{ kA}$
  - increase repetition rate (1.35 ightarrow 1.16 sec/cycle)
  - $\sim$  10% increased  $\nu$  flux at HK

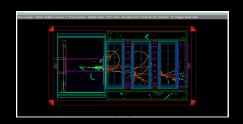


## near detector complex (280 m downstream)

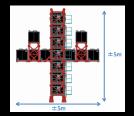
- monitor flux before oscillations
- measure cross-sections
- essential to tune fits and reduce systematics
- ▶ interaction dominated by CCQE
  - → single ring
  - $\rightarrow$  kinematic reconstruction of  $E_{\nu}$
- interaction rates uncertainty 3%
  - $\rightarrow$  systematic errors 5-6% in T2K
  - must be reduced to 3% for HK

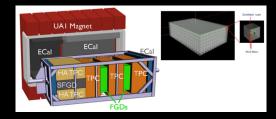


## near detector complex (280 m downstream)

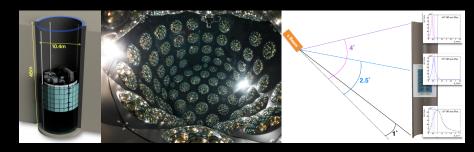


- on-axis beam monitor (Fe + scintillator),
   0.3 mrad already achieved in T2K
- off-axis magnetized detectors to identify wrong-sign bkg
- 2024 upgrades: TOF detectors, high-angle TPCs, scintillator superFGD
- improvements: finer granularity and wider solid angle





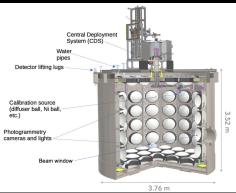
### intermediate water-Cerenkov detector (IWCD)

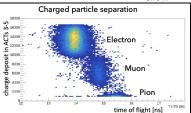


- ► 640 t water Cerenkov cylinder (like FD
- ▶ 368 ID 8 cm multiPMTs (+ 368 OD)
- ▶ 850 m downstream (reduce pile-up wrt ND)
- located on surface in 50 m pit
- movable to sample  $(1^{\circ}, 4^{\circ})$  off-axis angles  $(2.5^{\circ})$  in HK)
- match the flux between near and far detectors
- bdirect measurement of  $\nu_e/\nu_\mu$  cross-section as a function of energy (main systematic in  $\delta_{CP}$ )
- **measure** intrinsic  $\nu_e$  and NC components (more abundand off-axis)



### prototype tested @ CERN

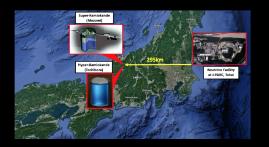




- IWCD half-sized prototype (same photosensors, calibration and water system)
- data taking 2024-2025 (including Gd phase)
- operated in CERN T9 (East Hall) beam line
- ▶ particles:  $e^{\pm}, \mu^{\pm}, \pi^{\pm}, p^{+}$ , tagged  $\gamma$
- ightharpoonup momentum:  $0.2-1.1~{\rm GeV/c}$
- beam detectors provide PID, calorimetry and triggering



## Hyper-Kamiokande far detector site



same choices as T2K:

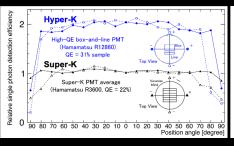
- 295 km baseline
- $ightharpoonup 2.5^{\circ}$  off-axis angle

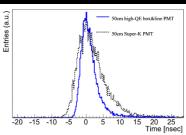
HK will be started while SK keeps running

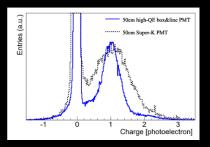




### ID photosensors: 50 cm PMTs



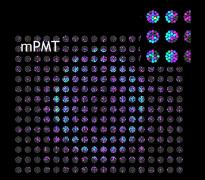




- 20k 50 cm Hamamatsu B&L PMT
- wrt SK: 2× single p.e. efficiency
- ➤ 2× increased charge (35%, 50% in SK) and time resolution (1.1 ns transit time spread, 2.1 ns in SK)
- 2× increased pressure resistance (70 m water depth)
- 6 kHz dark rate
- pressure-resistant acrylic cover

### ID photosensors: 800 mPMTs





- module with 19 3" PMTs
- high granularity and good time resolution (0.6 ns TTS,  $2 \times$  better than B&L)
- directional sensitivity from local coincidences
- small magnetic field sensitivity
- combined with 50 cm PMTs in HK and adopted also in IWCD

#### PMTs status

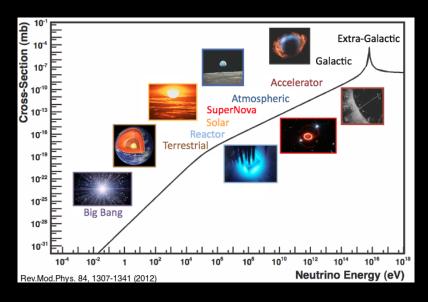
- mass production of 50 cm PMTs ongoing
- >15k PMTs have been delivered
- screen testing signal and visual inspection ongoing at Kamioka and Hamamatsu
- testing of covers using mock-up frame in Mediterranean sea



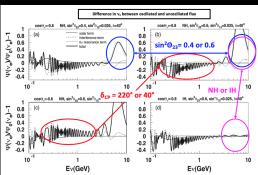




## very wide physics program



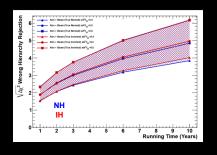
### atmospheric neutrinos

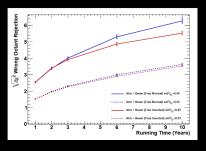


- wide energy range (peak 1 GeV) and long travel distance
- precision mixing parameters
- >  $\nu_e$  flux enhanced depending on mass ordering,  $\theta_{23}$  and  $\delta_{\mathrm{CP}}$



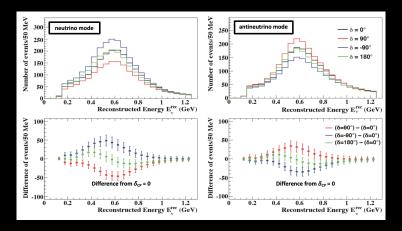
## mass ordering and octant sensitivity





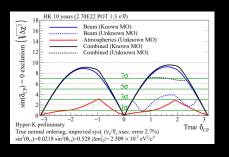
- ightharpoonup earth matter effect for u's passing through core allow measurement of  $\delta_{\rm CP}$ , mass ordering and  $\theta_{23}$  octant
- lacktriangle measure mass ordering at  $3\sigma$  for all values of  $heta_{23}$
- determine octant at  $3\sigma$  for  $|\theta_{23} 45^{\circ}| \geq 2^{\circ}$

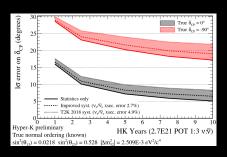
#### beam: CP violation



- $ightharpoonup 
  u_e$  appearance in  $u_\mu$  beam
- ► T2K disfavour CP conservation ( $\delta = 0, \pi$ ) at  $\sim 3\sigma$  level
- ightharpoonup effect of  $\delta_{CP}$  clearly seen in reconstructed spectra (10 years, NH)

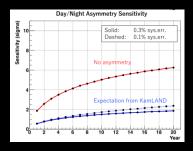
#### beam: CP violation

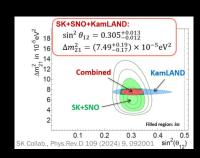


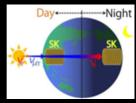


- ▶ assume 10 years exposure, 1:3  $\nu$  to  $\bar{\nu}$  beam operation, known mass ordering, 3-4% systematic uncertainty
- ightharpoonup exclude  $\sin\delta_{\sf CP}=0$  at  $5\sigma$  for 60% of range
- lacktriangle measure  $\delta_{\sf CP}$  with  $\pm (7^\circ, 22^\circ)$  for  $\sin \delta_{\sf CP} = (0, 1)$
- sensitivity paper: https://arxiv.org/abs/2505.15019

#### solar neutrinos

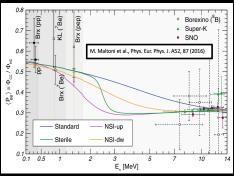


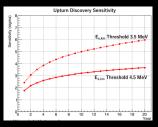




- ▶ regeneration of  $\nu_e$  due to earth matter effect  $\rightarrow$  higher  $\nu_e$  flux at night
- $\blacktriangleright$  can reject no day/night asymmetry to  $5\sigma$  after 10 years
- $\blacktriangleright$  also, resolve  $1.5\sigma$  tension between solar experiments and KamLAND on  $\Delta m^2_{21}$

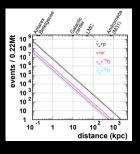
### solar neutrinos



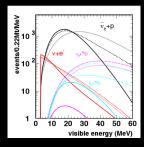


- solar neutrinos up-turn: increase in low-energy survival probability
- $>3-5\sigma$  measurement of up-turn spectrum after 10 years

## supernovae neutrinos: $\bar{\nu}_e + p^+ \rightarrow e^+ + n$

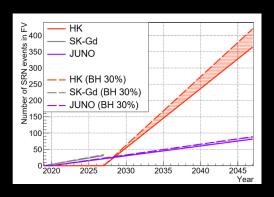


SN location	n events/tank		
galactic centre (10 kpc)	70k		
SN1987a (50 kpc)	3k		
Andromeda (0.8 Mpc)	10		



- first and only detection of supernova neutrinos SN1987a (11  $\pm$  8 events @ 50 kpc)
- ► 50-80k events for a 10 kpc supernova → neutrino properties, SN mechanism
- sensitive to 1 Mpc (Andromeda)
- ▶ 1.3° pointing for SN alert

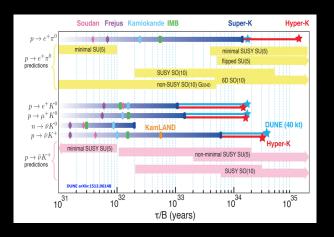
### diffuse supernova neutrino background



- neutrinos from all SN since the beginning of the universe
- relevant for stellar formation rate, heavy elements synthesis, black-hole formation
- lacktriangle expect 4 events/year with neutron tagging,  $4\sigma$  discovery after 10 years
- ▶ 16-30 MeV window limited by spallation backgrounds

also: dark matter annihilation in the sun  $\rightarrow$  neutrinos!

### proton decay

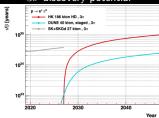


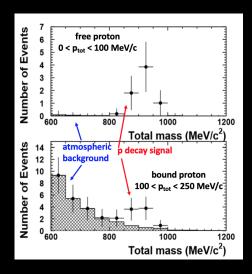
- predicted by many GUT models
- water Cerenkov sensitive to several decay modes
- suppress backgrounds with improved neutron tagging
- ightharpoonup reach  $10^{35}$  y lifetime

## proton decay: $p^+ o e^+ + \pi^0$

#### golden channel

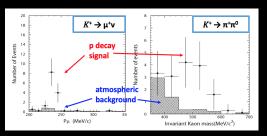
- favoured by many GUTs
- ► HK can see all products
- background-free search thanks to neutron tagging  $(n+p^+ \rightarrow d + \gamma, E_{\gamma} = 2.2 \text{ MeV})$
- $ightharpoonup 3\sigma$  discovery potential

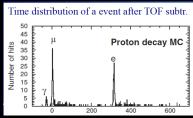


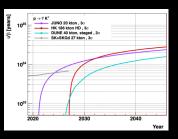


## proton decay: $p^+ \rightarrow \bar{\nu} + K^+$

- ►  $K^+$  below Cerenkov threshold (visible in DUNE), detect via decay  $K^+ \rightarrow \mu^+ + \nu$  (64%) or  $K^+ \rightarrow \pi^+ + \pi^0$  (21%)
- prompt gamma de-excitation, followed by  $\mu^+$  from  $K^+$  decay (12 ns) and  $e^+$  from  $\mu^+$  decay (2 ms)
- $ightharpoonup 3\sigma$  discovery



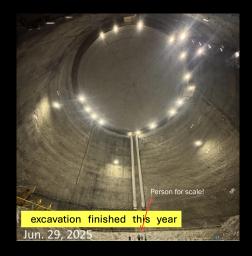




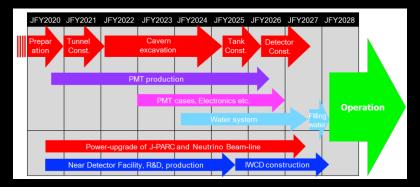
many other modes improved by 1 order

### excavation status





#### timeline



- design report released 2018
- cavern excavation finished 2025
- water filling 2027
- data taking 2028

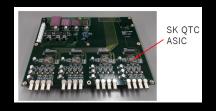
#### summary

- ► Hyper-Kamiokande is the next-generation neutrino experiment
  - ► 8× larger than SK
  - improved photo-sensors
  - upgraded 1.3 MW beam
  - ▶ intermediate water Cerenkov to reduce systematic errors
- rich physics program
  - precise neutrino oscillation:  $\delta_{CP}$ , mass ordering
  - astrophysics: solar neutrinos, supernovae etc.
  - new physics: proton decay, dark matter ...
- ► data taking will start in 2028

# Back-up slides

#### **Electronics**

- electronics timing resolution < PMT TTS</p>
- frontend electronics and HV module in water
- timing delivered by GPS
- communication via fiber (200 m)
- large buffer for SN burst
- prototype developed with: SK QTC ASIC (charge to time), FPGA TDC 4GHz
- electronics resolution: 0.2 ns at 1 p.e., 10% charge resolution at 1 p.e., 2000 p.e. range with 1% linearity



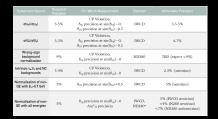
#### mPMT electronics:

- 0.4 ns for 1 p.e., 0.1 ns for larger pulses
- charge resolution: 0.05 to 25 p.e.
- power consumption: 3-4 W/mPMT for HK (driven by water circulation requirements), 5-10 W/mPMT for IWCD

### Systematic errors

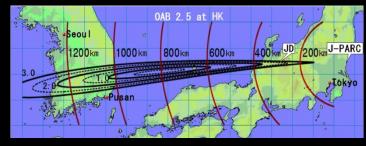
- statistical error will be 3% in HK
- require 3% systematic error on interaction rates (through upgraded ND280 and IWCD)

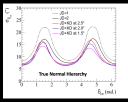
Source of uncertainty	$\nu_{\mu}$ CC	$\nu_{\rm e}$ CC	$\bar{\nu_{\mu}}$ CC	$\bar{\nu}_{\rm e}$ CC
Flux and common cross sections				
(w/o ND280 constraint)	10.8%	10.9%	11.9%	12.4%
(w/ ND280 constraint)	2.8%	2.9%	3.3%	3.2%
Unconstrained cross sections	0.8%	3.0%	0.8%	3.3%
SK	3.9%	2.4%	3.3%	3.1%
FSI + SI(+ PN)	1.5%	2.5%	2.1%	2.5%
Total				
(w/o ND280 constraint)	11.9%	12.2%	13.0%	13.4%
(w/ ND280 constraint)	5.1%	5.4%	5.2%	6.2%

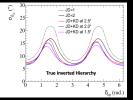


Systematic Source	Required Precision	For Which Measurement	Detector	Achievable Precision
Beam Direction	0.6 mrad (4 MeV shift)	$\delta_{op}$ precision at $sin(\delta_{op})\!\sim\!0$ $\Delta m^2_{32}$ precision	INGRID	<0.3 mrad (<2 MeV)
Removal (binding) energy	4 MeV*	$\delta_{cp}$ precision at $\sin(\delta_{cp}) - 0$ $\Delta m^2_{32}$ precision	IWCD, ND280	2.6 MeV (IWCD on O) -1 MeV (ND280 on C)**
High angle measurement (cosθ<0.2)	4%	CP Violation, $\delta_{\rm sp} \ precision \ at \ sin(\delta_{\rm sp}){\sim}0$	IWCD, ND280	<4% statistical precision in both detectors
Beam rate monitoring	~1% per day	General monitoring of beam quality	INGRID	<0.5% per day for neutrinos and antineutrinos
Neutron Multiplicity	TBD	Atmospheric neutrino Nucleon decay	IWCD, ND280	<5% IWCD <4% ND280
μπ <sup>0</sup> cross section & neutron multiplicity	TBD	en <sup>0</sup> proton decay	IWCD	TBD

#### Second tank in Korea

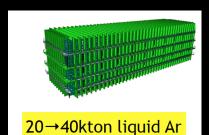






- 1100 km
- 2nd oscillation maximum
  - higher precision on  $\delta_{\mathrm{CP}}$

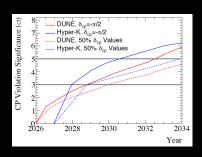
## HyperK and DUNE

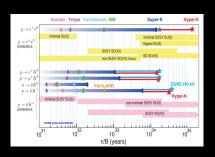


- HK: 188 kton water, 1.3 MW beam
- ▶ DUNE:  $20 \rightarrow 40$  kton liquid Ar,  $1.2 \rightarrow 2.4$  MW beam



## HyperK and DUNE





- ightharpoonup similar sensitivities for  $\delta_{\mathsf{CP}}$
- similar sensitivities for proton decay
  - DUNE has excellent event reconstruction, high efficiency, low backgrounds, golden  $p^+ \rightarrow \bar{\nu} + K^+$
  - $\blacktriangleright$  HK has larger volume and covers several decay modes, golden  $p^+ \to e^+ + \pi^0$
- supernovae
  - **DUNE** observes  $\nu_e \rightarrow$  time and flavour profile of collapse
  - ► HK has larger mass (188 kton vs 20/40 kton fiducial mass)

## Hyper-Kamiokande collaboration

