Neutrino Interactions from GeV to TeV Scale

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CPNR Workshop, October 24-27, 2025 Chonnam National University

My experience with neutrinos

The first Lecture

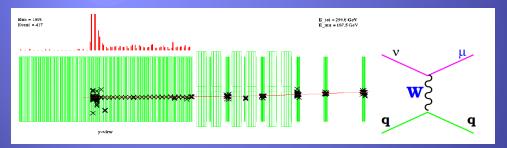
- PhD experiment with CCFR/NuTeV at Fermilab (1993-2000)
 - Differential Cross Sections and Structure Function Analyses (PRL 87, 251802 (2001), PRL 87, 251802 (2001))
 - Global Structure Functions analyses (PRL 82, 2467 (1999))
 - Phenomenology studies (NNLO, non-pQCD effect etc): (Eur Phys J C13, 241 (2000))
- Postdoc with CDF at Fermilab (2000~2005)
 - Modelling Neutrino Cross Section: J. Phys. G 29 (2003) 1899,
 while performing main job, CDF top quark physics

The Second Lecture

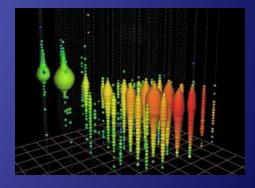
- > ATLAS / CMS at Manchester (2006~2012) / Seoul National Univ. (2013~)
 - Search for heavy neutrinos

Neutrino Interactions

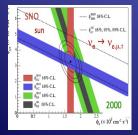
Neutrino interactions at high Energy (10~500 GeV) have been used to probe parton dynamics inside nucleons for precision QCD studies, and electroweak physics in 1980~2000 (CDHSW, CHARM, CHORUS, CCFR, NuTeV)



Ultra high energy neutrino interactions(>TeV scale) for astrophysics (IceCube, Antares)



Neutrino Oscillations brings a precision QCD studies back, but in Ev = few GeV region (T2K, Nova, Hyper-K, DUNE etc) since 2000







Precision QCD in neutrino

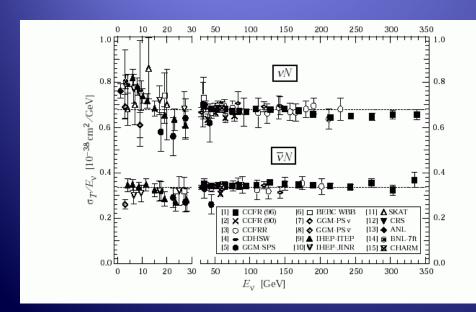
High energy neutrino beam (10 ~ 300 GeV)

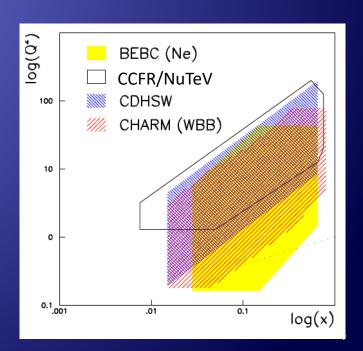
$$\frac{d^2\sigma}{dxdy}^{\nu(\overline{\nu})} \ = \ \frac{G^2ME}{\pi} \left[(1-y-\frac{Mxy}{2E})F_2 + \frac{y^2}{2}2xF_1 \pm y(1-\frac{y}{2})xF_3 \right]$$

$$F_{2,LO} = \sum_{i=u,d..} xq_i(x,Q^2) + x\overline{q_i}(x,Q^2) \quad xF_{3,LO} = \sum_{i=u,d..} xq_i(x,Q^2) - x\overline{q_i}(x,Q^2)$$

$$xF_{3,LO} = \sum_{i=u,d...} xq_i(x,Q^2) - x\overline{q_i}(x,Q^2)$$

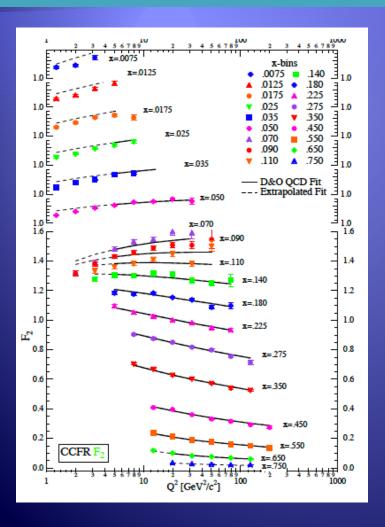
$$R = \frac{F_L}{2xF_1} = \frac{F_2}{2xF_1}(1 + Q^2/\nu^2) - 1.$$

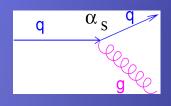


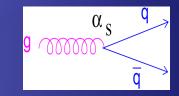


Precision QCD using Structure Functions (SFs)

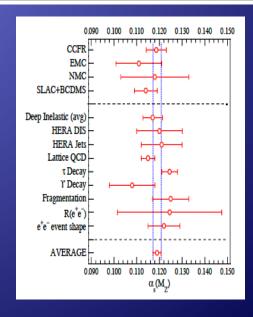
ightharpoonup DIS region: structure function QCD analysis (F₂, xF₃) With DGLAP evolution from high Q² to low Q²: extract $\Lambda_{\rm QCD}$ & PDFs



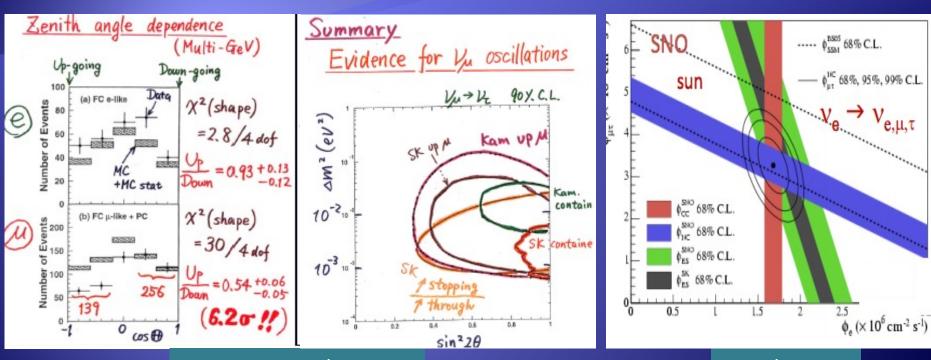




$$\frac{\partial q(x,Q^2)}{\partial \ln Q^2} = \frac{\alpha_s}{2\pi} \int_x^1 \frac{dy}{y} \left(q(y,Q^2) P_{q \to qg} \left(\frac{x}{y} \right) + g(y,Q^2) P_{g \to q\bar{q}} \left(\frac{x}{y} \right) \right)$$



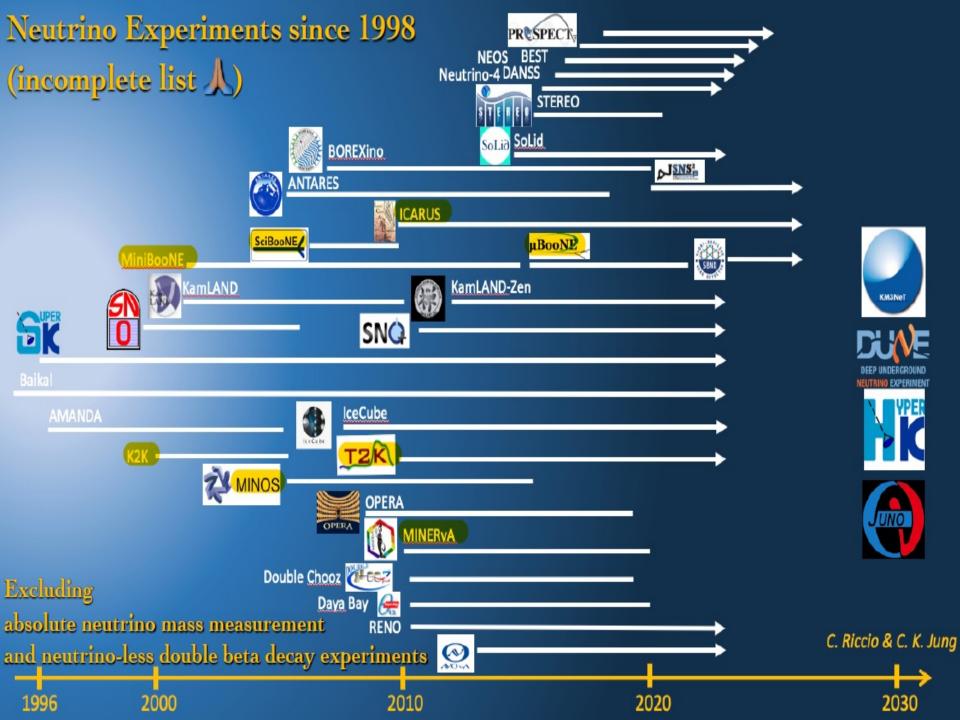
Neutrino Oscillations discovered



Neutrino 1998 by Kajita

2002 by SNO

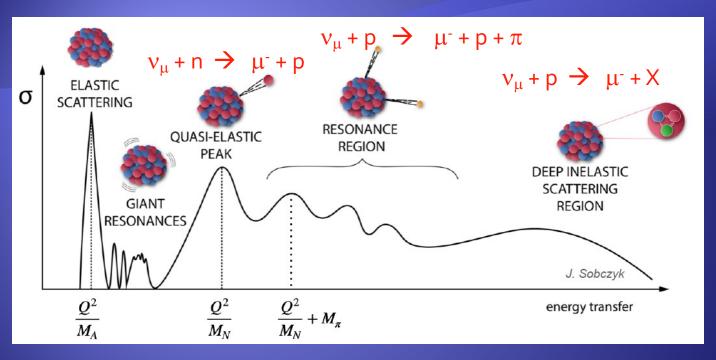
 \triangleright Discovery of Neutrino Oscillations in 1998 brings precision neutrino physics in QCD, but in Ev = few GeV region

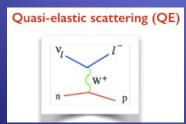


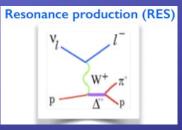
Physics Goals with Neutrinos

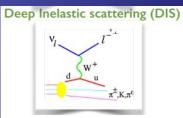
- Understanding of the matter-antimatter asymmetry in the Universe
 - CP violation
 - Leptogenis
- Source of non-zero neutrino mass
 - Seesaw mechanism
 - BSM model for Grand Unification
 - Dirac vs Majorana
- Absolute scale of neutrino mass
- Study of astrophysical high E neutrinos
- Precision measurements for new physics
 - Precision measurements of neutrino oscillation parameters (PMNS matrix)
 - Non-standard interactions

Neutrino-Nuclei Interactions





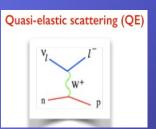


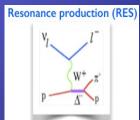


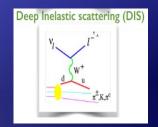
- Neutrino oscillation experiments: beam energies from 0.1 to 10 GeV
- > Resonance region overlapped with DIS region

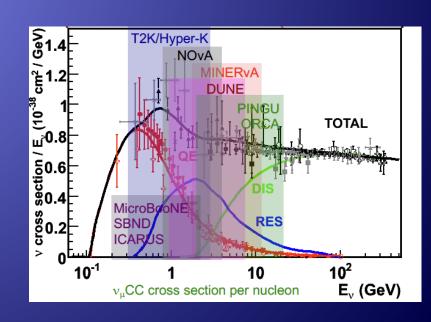
Neutrino-Nuclei Interactions

- Quasi-Elastic / Elastic (W=Mn): $v_u + n \rightarrow \mu^- + p$
- well measured and described by form factors
- Resonance (low Q², W< 2 GeV): $v_{\mu} + p \rightarrow \mu^{-} + p + \pi$
- poorly measured, only 1st resonance described by Regin and Seghal model
- □ Deep Inelastic (high Q2) $\nu_{\mu} + p \rightarrow \mu^{-} + X$
- well measured by high energy experiments and well described by quark-parton model, but not at low Q² region
- Large overlap with resonance region







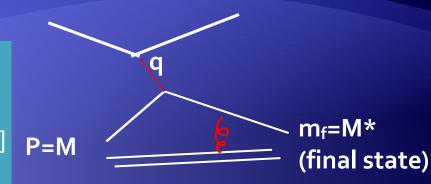


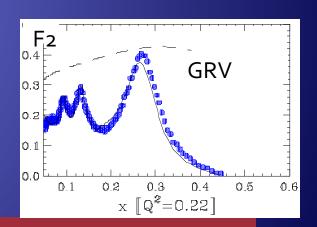
Neutrino-Nuclei Cross Sections

Can we build up a model to describe all Q² region from high down to very low energies ?

[DIS, resonance, even down to Q²=0 limit]

- Describe them in terms of quark-parton model including resonance scattering
- With PDFS, it is straightforward to convert charged-lepton scattering cross sections into neutrino cross section.
 (just matter of different couplings)





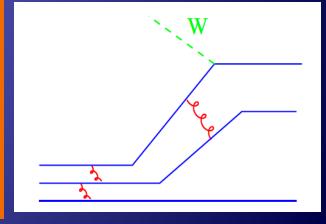
- Major Challenges
 - non-perturbative QCD effects (not resolving a single quark)
 - Large uncertainty on PDFs at high x & low Q²
 - Overlapped with a DIS contribution
 - Additional axial-vector contribution in neutrino scattering

Non-pQCD effect at low Q²

- Higher twist effect (dynamic): at low Q², not able to resolve a single parton, but scattering with multi-partons
- The contributions from higher twist diagrams are suppressed by powers of 1/Q²
- > Empirical model: use data to fit the higher twist effect using an empirical formular.

$$F_2(LT) (1+h2(x)/Q^2 + h4(x)/Q^4)$$

- Renormalon model: predict Q², and x dependence on SFs: get the overall size from the data
- Enough SLAC/Jefferson Lab e-N DIS data in low Q² region not used by PDF groups



$$F_2(x, Q^2) = F_2^{pQCD+TM} \left[1 + \frac{D_2(x, Q^2)}{Q^2} + \frac{D_4(x, Q^2)}{Q^4} \right]$$

$$D_2(x, Q^2) = \frac{a_2}{q(x, Q^2)} \int_x^1 \frac{dz}{z} c_2(z) q(x/z, Q^2)$$

Non-pQCD effect at low Q²

- Kinematic higher twist (target mass effect)
 - Nucleon target mass effect is not negligible
 - → use Nachtmann variable

$$\xi = \frac{2x}{1 + \sqrt{1 + 4x^2M^2/Q^2}} \ .$$

--> Full calculations

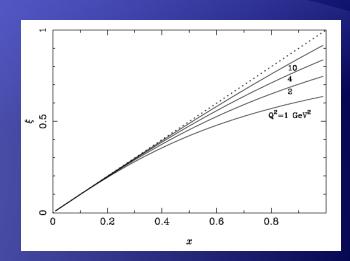
$$F_2^{\text{TMC}}(x, Q^2) = \frac{x^2}{\xi^2 r^3} F_2^{(0)}(\xi) + \frac{6M^2 x^3}{Q^2 r^4} h_2(\xi) + \frac{12M^4 x^4}{Q^4 r^5} g_2(\xi)$$

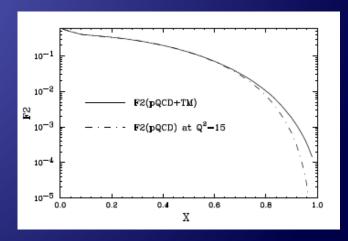
$$h_2(\xi, Q^2) = \int_{\xi}^1 du \, \frac{F_2^{(0)}(u, Q^2)}{u^2}$$

$$h_2(\xi, Q^2) = \int_{\xi}^1 du \, \frac{F_2^{(0)}(u, Q^2)}{u^2} \qquad g_2(\xi, Q^2) = \int_{\xi}^1 du \, h_2(u, Q^2) = \int_{\xi}^1 du \, \int_u^1 dv \, \frac{F_2^{(0)}(v, Q^2)}{v^2}$$

$$r = \sqrt{1 + \frac{4x^2M^2}{Q^2}}$$

A dominant player: using Nachtman variable

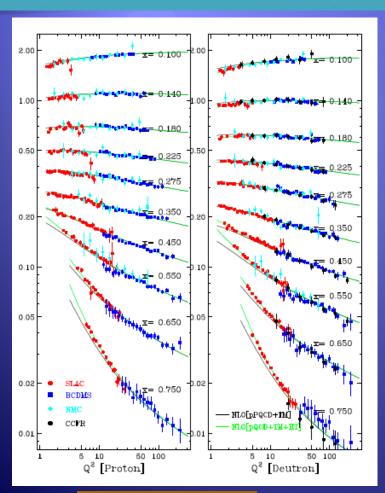


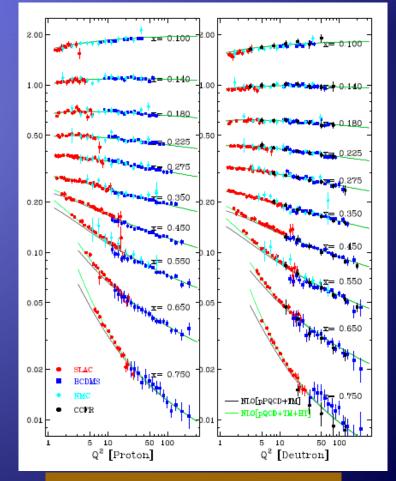


Higher Twist Effects 1

Both higher twists models describe the SFs data (0.1<x0.75)</p>

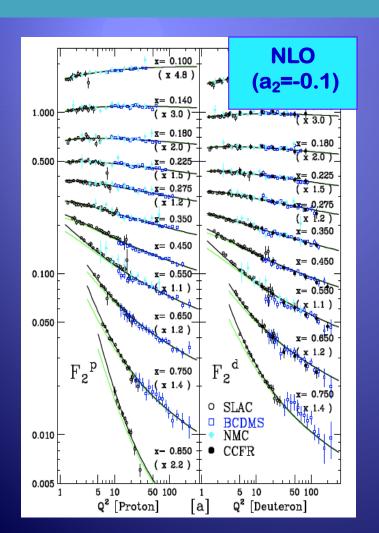
(SLAC, BCDMS, NMC, and R): CCFR shown for comparison

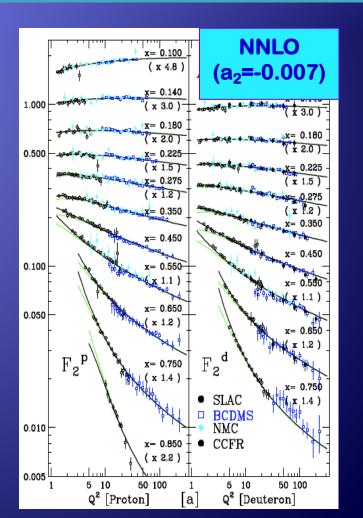




Higher Twist Effects 2

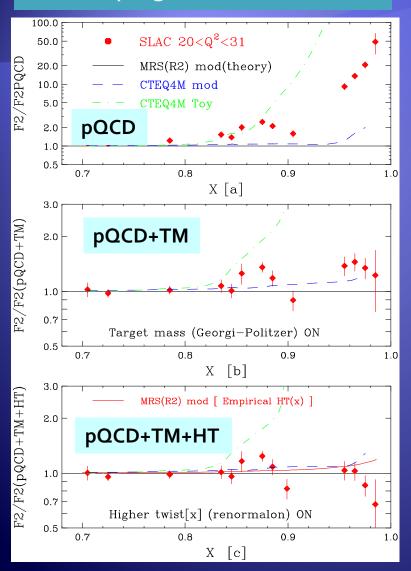
- Higher twist effects in NLO are similar to missing NNLO terms
- > NNLO pQCD+TM can describe the non-perturbative QCD effects at low Q²





Higher Twist Effects 3

Very high x and low Q²



Parton Distributions, d/u, and Higher Twist Effects at High x

U. K. Yang and A. Bodek Phys. Rev. Lett. **82**, 2467 – Published 22 March 1999

- Very high x and low Q² data is well described by the pQCD+TM+HT
- Extraction of the high x PDF is promising (1999)
 - still a large uncertainty (2025)
 at high x

Lessons from QCD studies

- NLO & NNLO analyses with DIS data: PRL 82, 2467 (1999), Eur. Phys. J. C13, 241 (2000) by Bodek and Yang
 - Kinematic higher twist (target mass) effects are large
 - Resonance region can be well described (duality works)
 - Most of dynamic higher twist corrections (in NLO analysis) are similar to missing NNLO higher order terms
 - Valence quarks PDFs up to x=0.95 in good agreement with data
- NNLO pQCD+TM with NNLO PDFs can describe the nonperturbative QCD effects at low Q²
- > Thus, we reverse the approach to build the model:
 - Use effective LO PDFs with a news scaling variable, ξw to absorb target mass, higher twist, missing QCD higher order

$$x_{Bj} = \frac{Q^2}{2M\nu} \Longrightarrow \xi_W = \frac{Q^2 + B}{\{M\nu[1 + \sqrt{(1 + Q^2 / \nu^2)}] + A\}}$$

$$F_2(x,Q^2) \to \frac{Q^2}{Q^2 + C} F_2(\xi_w,Q^2)$$

Modeling Cross Sections for all Q2 regions

- For MC Generator, Bodek-Yang LO approach: (pseudo NNLO)
 - Use effective LO PDFs with a new scaling variable, ξw to absorb target mass, higher twist, missing QCD higher orders

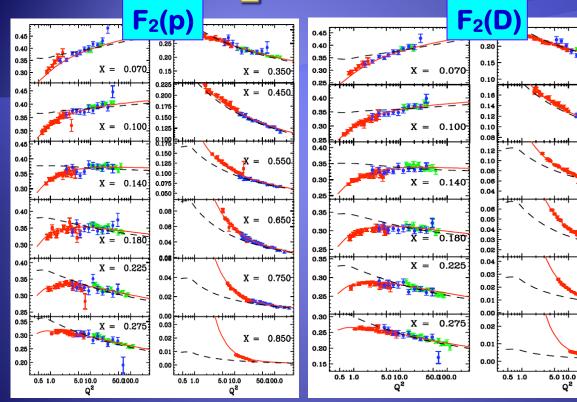
$$\xi_{W} = \frac{Q^{2}}{2Mv} \longrightarrow \xi_{W} = \frac{Q^{2} + B}{\{Mv[1 + \sqrt{(1 + Q^{2}/v^{2})}] + A\}}$$

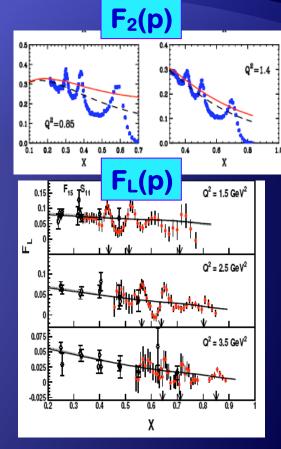
 Multiply all PDFs by K factors for photo production limit and higher twist

$$F_2(x,Q^2) \to \frac{Q^2}{Q^2 + C} F_2(\xi_w,Q^2)$$

- Kval (u,d) = $[1-G_D^2(Q^2)] * [Q^2+C_{2V}] / [Q^2+C_{1V}]$, where $G_D^2(Q^2) = 1/[1+Q^2/0.71]^4$ (motivated by Adler Sum Rule)
- Ksea $(u,d,s) = Q^2/[Q^2+Csea]$

Fit to F₂ data and Predictions





- Effective LO
- --- X_{bj}

Excellent Fitting results on all DIS $F_{2,}(\chi^2/DOF=1235/1200)$

X = 0.750

X = 0.850

50.000.0

- Predictions for resonance region F₂, F_L
- Quark-hadron duality works

Photo-production data

In the BY model, freeze the DGLAP evolution at Q² = Q²_{min}

$$F_2(x, Q^2 < 0.8) = K(Q^2) * F_2(\xi w, Q^2 = 0.8)$$

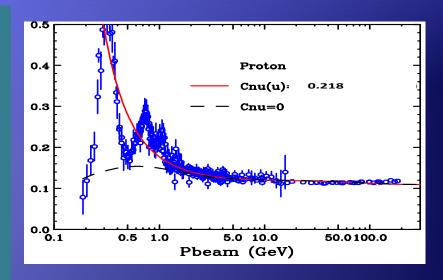
Additional K^{LW} factor for valence quarks:

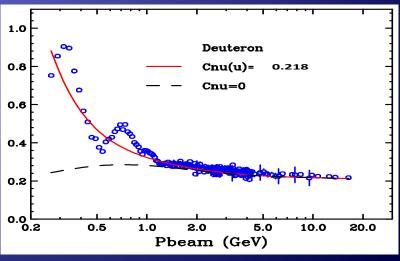
Kval =
$$\frac{K^{LW}*[1-G_D^2(Q^2)]}{*[Q^2+C_{1V}]}$$

$$K^{LW} = (v^2 + C^v)/v^2$$

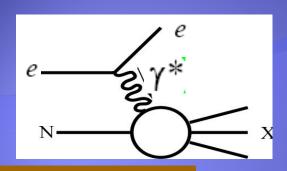
This makes a duality work all the way down to Q²=0 (for charged leptons)

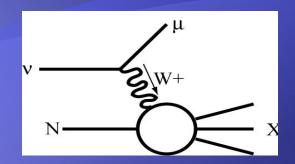
> Photo-production data with v(Pbeam)>1 GeV included in the fitting





e-N vs v-N scattering cross sections





e-N DIS scattering

$$\frac{d^2\sigma}{dxdy} = \frac{4\alpha^2 ME}{Q^4} \left[\frac{y^2}{2} 2x F_1(x, Q^2) + (1 - y) F_2(x, Q^2) \right]$$

ν-N DIS scattering

$$\frac{d^2 \sigma^{v,\bar{v}}}{dxdy} = \frac{G_F^2 ME}{\pi} \left[\frac{y^2}{2} 2x F_1(x, Q^2) + (1 - y) F_2(x, Q^2) \pm (y - \frac{y^2}{2}) x F_3(x, Q^2) \right]$$

■Charged-lepton (e, μ) scattering

F2(p) = x[4/9(u+u) + 1/9(d+d) + 1/9(s+s)]
F2(n) = x[4/9(d+d) + 1/9(u+u) + 1/9(s+s)]
assuming
$$d_p = u_n$$

Neutrino scattering

$$F_2(v) = x[(u+u) + (d+d) + (s+s)]$$
$$xF_3(v) = x[(u-u) + (d-d)]$$

Neutrino cross sections

- Figure 10 model using parton model with ξw describe all DIS and resonance F_2 data as well as photo-production data ($Q^2=0$ limit): vector contribution works well
- Neutrino Scattering:
 - Do we see an axial vector contribution at low Q²?
 - Nuclear effect in neutrino? different from charged lepton scattering data?
 - Heavy charm mass effect ? (W*+s \rightarrow c) different from charged lepton scattering (g \rightarrow cc, $\gamma*$ +c \rightarrow c)
 - use ξ w slow rescaling algorithm for F_2 , $2xF_1$, and xF_3
 - Effective LO model works for xF₃? Yes with NLO correction f(x)
 - Use R=R₁₉₉₈ to get 2xF₁

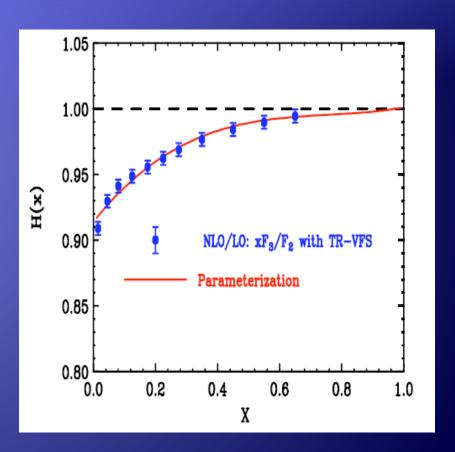
Effective LO model for xF₃?

- Scaling variable, ξw absorbs higher order effect for F_{2,} but the higher order effects for xF₃ can be different
- Use NLO QCD to get double ratio

$$H(x) = \frac{xF_3(\text{NLO})}{xF_3(\text{LO})} / \frac{F_2(\text{NLO})}{F_2(\text{LO})}$$

not 1 but almost indep. of Q²

Enhance anti-neutrino cross section by 3%



Axial Vector Structure Functions

- At high Q², vector and axial vector contribution are same, but not at low Q²
- K factors for axial contributions: type II

$$K_{sea}^{vector} = \frac{Q^2}{Q^2 + C} \Rightarrow K_{sea}^{axial} = \frac{Q^2 + 0.55C_{sea}^{axial}}{Q^2 + C_{sea}^{axial}} \qquad K_{val}^{axial} = \frac{Q^2 + 0.1C_{val}^{axial}}{Q^2 + C_{val}^{axial}}$$

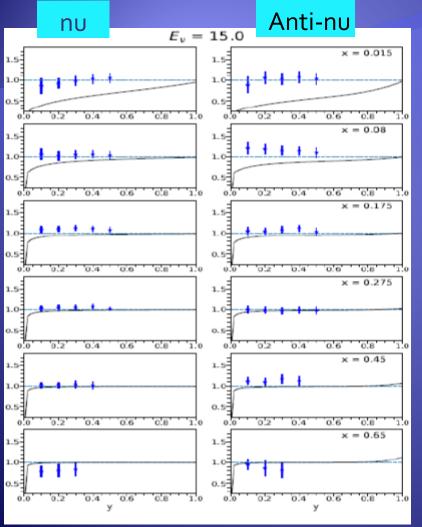
$$K_{val}^{axial} = \frac{Q^2 + 0.1C_{val}^{axial}}{Q^2 + C_{val}^{axial}}$$

where
$$C_{sea}^{axial} = 0.75$$
, $C_{val}^{axial} = 0.18$

- 0.55 was chosen to satisfy the prediction from PCAC by Kulagin, agrees with CCFR/CHROUS data for F_2 extrapolation to ($Q^2=0$)
- But, the non-zero PCAC component of F_2^{axial} at low Q^2 : mostly longitudinal

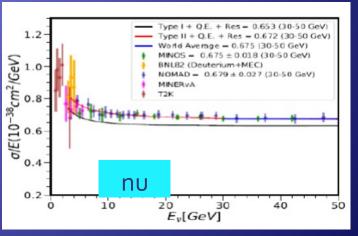
$$2xF_1^{axial} = 2xF_1^{vector}$$

Axial-vector contributions from neutrino data (CCFR (Fe), CHORUS (Pb))



Blue point: CHORUS/theory (type II)

- Solid line:theory (type I)/(type II)
- Type I (Vector = Axial at low Q²)
- Type II (Vector < Axial at low Q²)



→ Different axial vector contributions are clearly shown!

Bodek-Yang(BY) Model

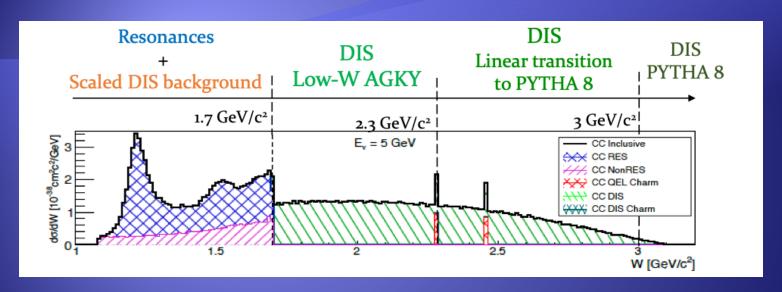
- BY Effective LO model with ξw describe all e/μ DIS and resonance data as well as photo-production data (down to Q²=0): provide a good reference for neutrino cross section
- Model also works well down to W=1.4 GeV, thus providing overlap with resonance models
- Low energy neutrino experiments can normalize their data to our model to extract their flux

J. Phys. G 29 (2003) 1899, arXiv.2108.09240

TABLE III. Summary of the neutrino event generators and choice of models used by each to generate the inclusive cross-section predictions against which comparisons to this measurement are made. RFG = relativistic Fermi gas, LFG = local Fermi gas, LS = Llewellyn Smith, RPA = random phase approximation, RS = Rein-Sehgal, BS = Berger-Sehgal, BY = Bodek-Yang, PY = PYTHIA, BUU = Boltzmann-Uehling-Uhlenbeck.

Generator	QE/MEC initial state	QE	MEC	Res	DIS	FSI
GENIE v2.12.2	RFG	L-S	Empirical	RS	BY+PY6	hA (data-driven empirical cascade)
GENIE v3.00.06	LFG	València	València	BS	BY + PY 6	hN (Oset (pions) + GENIE (nucleons))
NEUT v5.4.0	LFG	València	València	RS	BY + PY5	Oset + external data
NuWro 2019	LFG	L-S + RPA	València	BS	BY + PY 6	Oset $(pions) + NuWro (nucleons)$
GiBUU 2019	LFG	GiBUU Model			BY + PY 6	BUU Equations

Tuning resonance, inelastic scatting cross sections into MC generators



PRD 104 (2021) 072009

GENIE's Shallow-Inelastic Scattering model

RES

- Rein-Sehgal or Bergher-Sehgal are the starting point
- Added additional resonances
- Dipole Parameterization

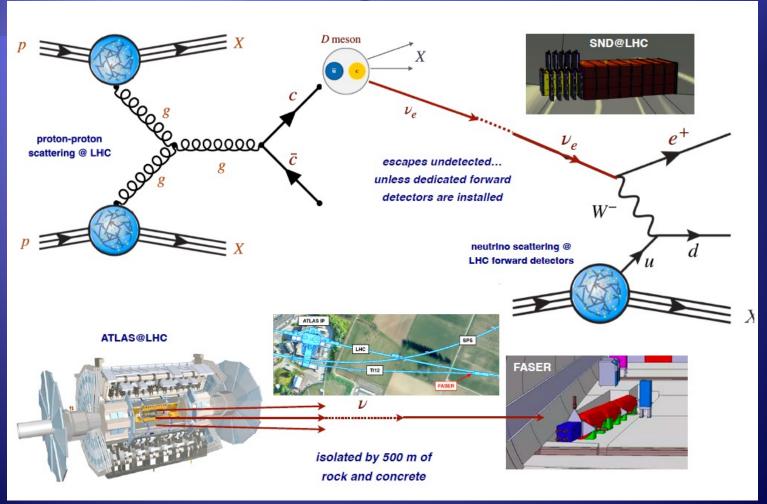
Non-resonant bkg

- · Duality-based approach
- Scaled Bodek-Yang model
- Scaling factors depend on initial state and hadron multiplicity
- Coupled to low-W AGKY model

DIS

- Bodek-Yang model
- Cross-section calculation at partonic level
- AGKY hadronization model

TeV neutrinos at the LHC: Faser experiment



Faser at the LHC:

Presented in ICHEP 2024

PHYSICAL REVIEW LETTERS 133, 021802 (2024)

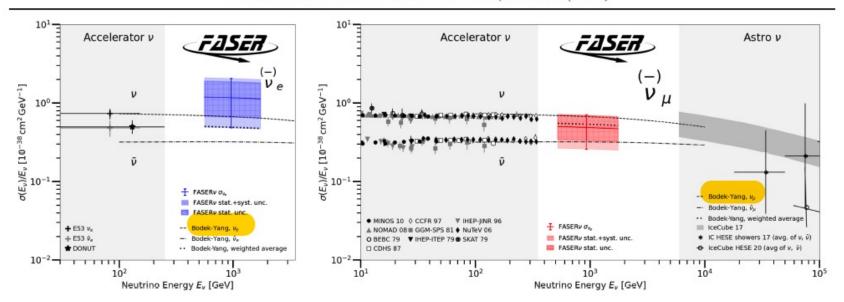
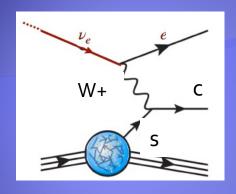
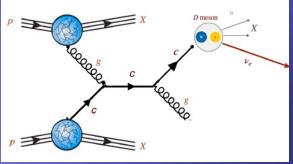
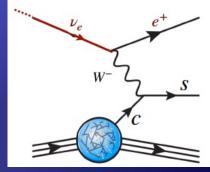


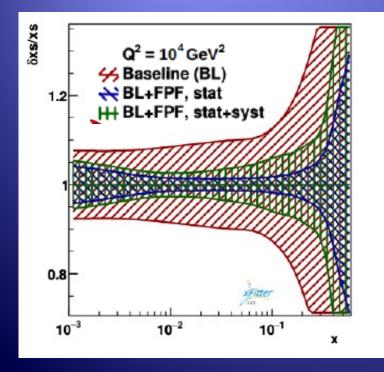
FIG. 4. The measured cross section per nucleon for ν_e (left) and ν_μ (right). The dashed contours labeled "Bodek-Yang" are cross sections predicted by the Bodek-Yang model, as implemented in GENIE. Note that the displayed experiments do not all use the same targets.

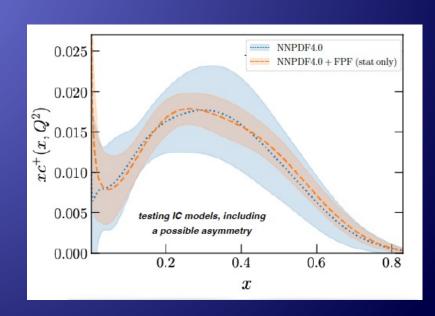
Impact of the LHC Neutrinos











Summary and Outlook

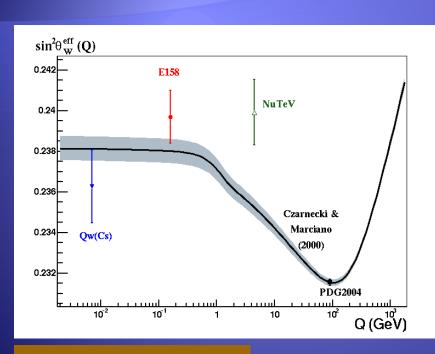
- Neutrino experiments have provided precision QCD studies using high energy and high intensity neutrino beams.
- Neutrino oscillation experiments brought new challenges in understanding neutrino interactions in low neutrino energy region
- Many improvements and important measurements have been made for last two decades in building models: resonance, SIS, DIS regions
- Bodek-Yang model have been used for DIS and DIS regions.
- For future neutrino oscillation experiments for DUNE, HyperK, it is important to understand following effects
 - Different nuclear effect in neutrino
 (LHC collider experiments, Drell-Yan, W, Z)
 - Axial vector contribution
 - Proper systematic treatments in old dataset is important
 - Inputs from LHC Forward Physics Programs will be useful



Thank You Bodek-Yang

Backup slides

Precision Neutrino data: electroweak mixing angle (CC / NC ratio)



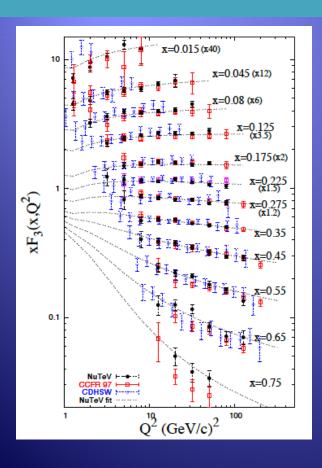
3σ effects bring many issues!

- Missing higher order QCD effect
- Different nuclear shadowing effect between NC and CC
- Charge symmetry violation
 (d quark pdf inside neutron
 vs u quark density inside proton)
- Asymmetry in strange sea (strange vs anti-strange quark)
- Possible to make an agreement within 1σ : PLB 693 (2010) 462 by Bentz et al.

PRL 88 (2002) 091802

Nuclear effect in neutrino?

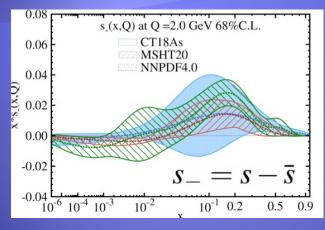
- Charged lepton scattering data (p,D) show good agreement with precision neutrino data in 0.01<x<0.6 and 1<Q²<100 GeV²</p>
- > But discrepancy shown in high x (>0.5) between CCFR and NuTeV data

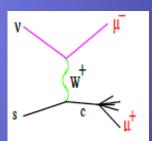


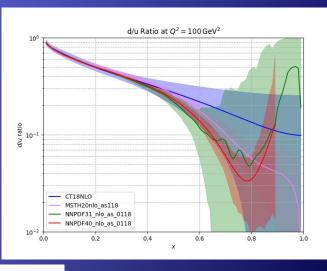
- NuTeV: refurbished CCFR detector,
 Better and precise calibration than
 CCFR calibrations
- Different neutrino effect at high x region?
- Future test will be necessary

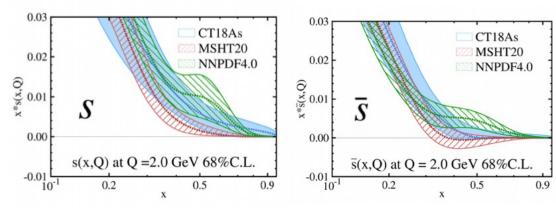
Asymmetry in strange quark

Tension from ATLAS W/Z data and neutrino dimuon data









Hou, Huey-Wen Lin, Mengshi Yan, C.-P. Yuan, 2211.11064]

DUNE Cross Section Modeling

Cross Section Modeling

DUNE TDR Cross Section Model (Eur. Phys. J. C 80, 978 (2020))

 $\pm 5\%$

- Example: DUNE TDR LBL Analysis
- All of the uncertainty in v-Ar modeling is captured in these ~30 parameters
 - The LBL analysis implicitly assumes that every other part of the xsec model is correct
 - e.g. the shapes of the distributions modified by these parameters are assumed to be correct; they need only be scaled by a correct parameter value
- The ND fit can strongly constrain these parameters
 - DUNE ND will collect nearly 1M events per week
- It turns out to be very easy to constrain xsec errors beyond the level at which we trust our models (see next slides)
 - In the DUNE TDR analysis, non-zero xsec errors were achieved by choosing minimum-allowed values for detector errors

Description	1σ
Quasielastic	
$M_{ m A}^{ m QE},$ Axial mass for CCQE	$^{+0.25}_{-0.15}~{ m GeV}$
QE FF, CCQE vector form factor shape	N/A
$p_{\rm F}$ Fermi surface momentum for Pauli blocking	$\pm 30\%$
Low W	
$M_{\Lambda}^{\mathrm{RES}}$, Axial mass for CC resonance	$\pm 0.05~{\rm GeV}$
$M_{ m V}^{ m RES}$ Vector mass for CC resonance	$\pm 10\%$
$\varDelta\text{-decay}$ ang., θ_π from \varDelta decay (isotropic \to R-S)	N/A
High W (BY model)	
$A_{\rm HT},$ higher-twist in scaling variable ξ_w	$\pm 25\%$
$B_{\Pi \Pi},$ higher-twist in scaling variable ξ_w	$\pm 25\%$
$C_{ m V1u},$ valence GRV98 PDF correction	$\pm 30\%$
C_{V2u} , valence GRV98 PDF correction	$\pm 40\%$
Other neutral current	
$M_{ m A}^{ m NCRES},$ Axial mass for NC resonance	$\pm 10\%$
Morning	

 $M_{\rm V}^{\rm NCRES}$. Vector mass for NC resonance

GENIE Xsec Parameters

	CEITIE I OI I GIGINOTOIO						
	Description	1σ					
_	N. CEX, Nucleon charge exchange probability	±50%					
=	N. EL, Nucleon elastic reaction probability	±30%					
	N. INEL, Nucleon inelastic reaction probability	$\pm 40\%$					
	N. ABS, Nucleon absorption probability	$\pm 20\%$					
	N. PROD, Nucleon π -production probability	$\pm 20\%$					
	π CEX, π charge exchange probability	$\pm 50\%$					
	π EL, π elastic reaction probability	$\pm 10\%$					
	π INEL, π inelastic reaction probability	$\pm 40\%$					
	π ABS, π absorption probability	$\pm 20\%$					
	π PROD, π $\pi\text{-production}$ probability	$\pm 20\%$					

GENIE FSI Parameters

Uncertainty	Mode
BeRPA [A,B,D]	$1p1h/\mathrm{QE}$
${\rm ArC}2p2h~[\nu,\bar{\nu}]$	2p2h
E_{2p2h} [A,B] $[\nu,\bar{\nu}]$	2p2h
NR $[\nu,\bar{\nu}]$ [CC,NC] $[n,p]$ $[1\pi,2\pi,3\pi]$	Non-res. pion
ν_e PS	$\nu_e, \overline{\nu}_e$ inclusive
$\nu_e/\overline{\nu}_e$ norm	$\nu_e, \overline{\nu}_e$ inclusive
NC norm	NC

Additional Xsec Parameters

Mike Wilking | The PRISM Technique for LBL v Experiments

NuINT, JGU Mainz, 2025/10/10